

1. (a) Given the Planck Distribution $\rho(\lambda)d\lambda = \frac{8\pi hc}{\lambda^5} \frac{1}{e^{hc/\lambda kT} - 1} d\lambda$

please derive Wein's displacement law: $\lambda_{\max} T = \text{constant}$ (10%)

$$c = \lambda\nu \rightarrow \nu = \frac{c}{\lambda} \Rightarrow d\nu = -\frac{c}{\lambda^2} d\lambda \quad (\text{turn into absolute value})$$

$$\text{Assume } x = \frac{hc}{\lambda kT}, dx = \frac{-hc}{\lambda^2 kT} d\lambda = \frac{-kT}{hc} x^2 d\lambda$$

$$\lambda_{\max} \text{ occurs when } \frac{d\rho(\lambda)}{d\lambda} = \frac{d\rho(\lambda)}{dx} \frac{dx}{d\lambda} = 0$$

$$\left(\frac{-kTx}{hc}\right)^2 \frac{d}{dx} \left[8\pi c \left(\frac{-kTx}{hc}\right)^5 \frac{1}{e^x - 1} \right] = 0$$

$$\left(\frac{-kTx}{hc}\right)^2 8\pi c \left(\frac{-kTx}{hc}\right)^5 \frac{d}{dx} \left[\frac{x^5}{e^x - 1} \right] = 0$$

$$\frac{d}{dx} \left[\frac{x^5}{e^x - 1} \right] = 0$$

$$\frac{5e^4(e^x - 1) - x^5 e^x}{(e^x - 1)^2} = 0$$

$$5(e^x - 1) - x e^x = 5 \left(1 - \frac{1}{e^x - 1} \right) = 0$$

By numerical method, $\Rightarrow x \approx 5$

$$\lambda = \frac{hc}{xkT} \Rightarrow \lambda_{\max} T \approx \frac{hc}{5k} = 2.9 \times 10^{-3} \text{ mK}$$

(b) Convert the expression in (a) to frequency domain. (5%)

$$c = \lambda\nu \Rightarrow \lambda = \frac{c}{\nu} \Rightarrow \frac{d\lambda}{d\nu} = -\frac{c}{\nu^2} \Rightarrow d\lambda = -\frac{c}{\nu^2} d\nu$$

$$\begin{aligned} R(\lambda) d\lambda &= \left(\frac{2\pi h}{\lambda^5}\right) \frac{c^2}{e^{\frac{hc}{\lambda kT}} - 1} d\lambda = \left(\frac{2\pi h}{\left(\frac{c}{\nu}\right)^5}\right) \frac{c^2}{e^{\frac{h\nu}{kT}} - 1} \times \frac{c}{\nu^2} d\nu \\ &= \left(\frac{8\pi h\nu^3}{c^2}\right) \frac{1}{e^{\frac{h\nu}{kT}} - 1} d\nu \end{aligned}$$

(c) What is Stefan-Boltzmann law? (5%) Please derive it from (a) or (b) (5%)

$$R = \sigma T^4$$

$$\text{Hint: } \int_0^\infty \frac{x^3}{e^x - 1} dx = \frac{\pi^4}{15}$$

$$E = \int_0^\infty \frac{8\pi hc}{\lambda^5} \frac{1}{e^{hc/\lambda kt} - 1} d\lambda = 8\pi hc \int_0^\infty \frac{1}{\lambda^3 \lambda^2} \frac{1}{e^{hc/\lambda kt} - 1} d\lambda$$

$$\text{let } u = \frac{hc}{\lambda kT}, \lambda = \frac{hc}{ukT}, (\lambda \rightarrow 0, u \rightarrow \infty)(\lambda \rightarrow \infty, u \rightarrow 0)$$

$$\begin{aligned} \frac{du}{d\lambda} &= \frac{hc}{kT}, \left(\frac{-1}{\lambda^2}\right), -\frac{kT}{hc} du = \frac{d\lambda}{\lambda^2} \\ &= 8\pi hc \int_{\infty}^0 \left(\frac{ukT}{hc}\right)^3 \frac{1}{(e^u - 1)} \frac{-kT}{hc} du = \frac{-8\pi k^4}{(hc)^3} T^4 \int_{\infty}^0 \frac{u^3}{(e^u - 1)} du \\ &= \frac{-8\pi k}{(hc)^3} T^4 \left[\frac{-\pi^4}{15}\right] = \left(\frac{8\pi^5 k^4}{15(hc)^3}\right) T^4 \\ &= \sigma T^4 \text{ (有寫出常數} \times T^4 \text{便給分)} \end{aligned}$$

(d) Sirius, one of the hottest known stars, has approximately a blackbody spectrum with $\lambda_{\max} = 260$ nm. From the Wein's displacement law $\lambda_{\max} T \cong hc/5k$, find the surface temperature of Sirius. (5%)

$$\lambda_{\max} T \approx \frac{hc}{5k} \rightarrow T = \frac{6.626 \times 10^{-34} \times 2.998 \times 10^{17}}{260 \times 10^{-9} \times 5 \times 1.381 \times 10^{-23}} = 11064K$$

2. For the ground-state of a harmonic oscillator, $\psi_0 = (\alpha/\pi)^{1/4} \exp(-\alpha x^2/2)$, $\alpha = 2\pi\nu m/\hbar$, calculate: (a) prove that ψ_0 is a solution of the Schrödinger equation with the eigenvalue $1/2 h\nu$. (10%)

薛丁格方程式：

$$\begin{aligned} &-\frac{\hbar^2}{2m} \frac{d^2\Psi(x)}{dx^2} + \frac{1}{2} kx\Psi(x) \\ \frac{d^2\phi_0}{dx^2} &= \left(\frac{\alpha}{\pi}\right)^{1/4} \left[\alpha e^{-\frac{\alpha x^2}{2}} - \alpha^2 x^2 e^{-\frac{\alpha x^2}{2}} \right] = (-\alpha + \alpha^2 x^2) \left(\frac{\alpha}{\pi}\right)^{1/4} e^{-\frac{\alpha x^2}{2}} \\ &= (-\alpha + \alpha^2 x^2) \phi_0 \\ &-\frac{\hbar^2}{2m} (\alpha^2 x^2 - \alpha) \phi_0 + \frac{1}{2} kx^2 \phi_0 = E \phi_0 \\ \alpha &= \frac{2\pi\nu m}{\hbar}, \nu = \frac{1}{2\pi} \sqrt{\frac{k}{m}} \rightarrow \alpha = \frac{\sqrt{mk}}{\hbar} \end{aligned}$$

代入得

$$-\frac{\hbar^2}{2m} \left(\frac{mk}{\hbar^2} x^2 - \frac{\sqrt{mk}}{\hbar} \right) + \frac{1}{2} kx^2 = E$$

化簡得

$$E = -\frac{1}{2} kx^2 + \frac{1}{2} h\nu + \frac{1}{2} kx^2 = \frac{1}{2} h\nu$$

- (b) For H₂ molecule, the vibrational frequency is 4403.2 cm⁻¹ what is the force constant (N/m) and the vibrational zero-point energy (in kcal/mol)? (10%)

$$\tilde{\nu} = \frac{\nu}{c}, \nu = \frac{1}{2\pi} \sqrt{\frac{k}{m}}$$

$$K = (443020 \times C \times 2\pi)^2 \times \frac{1.0078^2}{2 \times 1.0078} \times 1.66054 \times 10^{-27} \approx 576 \frac{N}{m}$$

$$E = \frac{1}{2} h\nu \approx 4.37 \times 10^{-20} J \approx 6.29 \text{ kcal/mol}$$

- (c) what are the vibrational frequencies of HD and D₂? (in cm⁻¹) (H = 1.0078, D =

2.0141 amu) (10%)

$$\tilde{\nu}(HD) = 4403.2 \times \sqrt{\frac{\left(\frac{1.0078 \times 1.0078}{1.0078 \times 2}\right)}{\left(\frac{1.0078 \times 2.0141}{1.0078 + 2.0141}\right)}} \left(\sqrt{\frac{\mu_{H_2}}{\mu_{HD}}}\right) \approx 3814 \text{ cm}^{-1}$$

$$\tilde{\nu}(D_2) = 4403.2 \times \sqrt{\frac{\left(\frac{1.0078 \times 1.0078}{1.0078 \times 2}\right)}{\left(\frac{2.0141 \times 2.0141}{2.0141 \times 2}\right)}} \left(\sqrt{\frac{\mu_{H_2}}{\mu_{D_2}}}\right) \approx 3113 \text{ cm}^{-1}$$

(d) if $\tilde{\nu}x_e = 121.3 \text{ cm}^{-1}$, what is the zero-point energy and the first excited-state energy? (5%)

$$E_0 = \frac{1}{2}h\nu - \frac{1}{4}h\nu x_e$$

$$E_v = \left(v + \frac{1}{2}\right)h\nu - (v + 1)^2 h\nu x_e, \quad \tilde{\nu} = 4403.2 \text{ cm}^{-1}$$

$$E_0 = \frac{1}{2}hc\tilde{\nu} - \frac{1}{4}hc(\tilde{\nu}x_e) = \frac{1}{2}hc\tilde{\nu} - \frac{1}{4}hc(12130) = 4.3 \times 10^{-20} \text{ (J)}, \quad 6.2 \text{ kcal/mol}$$

$$E_v = \left(v + \frac{1}{2}\right)h\nu - (v + 1)^2 h\nu x_e$$

$$= \left(v + \frac{1}{2}\right)hc\tilde{\nu} - (v + 1)^2 hc\tilde{\nu}x_e = 1.3 \times 10^{-19} \text{ (J)}, \quad 18.4 \text{ kcal/mol}$$

3. To a good approximation, the microwave spectrum of Al^{27}H consists of a series of equally spaced lines, separated by 12.604 cm^{-1} . Calculate the bond distance of AlH in \AA . (15%) ($\text{Al}^{27} = 26.9815 \text{ amu}$)

For microwave spectrum (rotational spectrum)

The energy difference between two states is $2B$, and for diatomic molecule

$$B = \frac{\hbar^2}{2I}, I = \mu R^2, E = hc \frac{1}{\lambda} = 2B$$

$$2B = hc(1260.4), 2 \times \frac{\hbar^2}{2I} = hc(1260.4)$$

$$I = \frac{h}{4\pi^2 c(1260.4)} = \mu R^2, \mu = \frac{1 \times 26.9815}{1 + 26.9815} \times 1.67 \times 10^{-27} \approx 1.61 \times 10^{-27}$$

$$R^2 = \frac{6.626 \times 10^{-34}}{4\pi^2 (3 \times 10^8)(1260.4)(1.61 \times 10^{-27})} \approx 2.76 \times 10^{-20}, \quad R \approx 1.67 \text{ \AA}$$

4. (a) Please justify the Dulong-Petit's law for the heat capacity of simple solids. (10%)

Dulong-Petit's law states that the molar heat capacity of solid elements is about $3R$.

Dulong and Petit proposed that the specific heat of all simple (monatomic) solids is equal, approximately 25 J/mol K . The average energy of each vibration would be kT , and the total vibrational energy of one mole of atoms would be $3RT$, leading to a specific heat of $3R$.

(Almost all experimental results before 1872 were consistent with this law.)

(b) The Einstein heat capacity formula are:

$$C_{v,m} = 3R f^2$$

$$f = \frac{\theta_E}{T} \left(\frac{e^{\theta_E/2T}}{e^{\theta_E/T} - 1} \right)$$

$$\theta_E = \frac{h\nu}{k}$$

Please show the above formula gives the classical result ($3R$) at high temperature and zero at very low temperature. (10%)

低溫下， T 趨近於 0 ，因為 $e^{\frac{\theta_E}{T}}$ 很大可以忽略 -1 ，

$$f \approx \frac{\theta_E}{T} \left(\frac{e^{\frac{\theta_E}{2T}}}{e^{\frac{\theta_E}{T}}} \right) \approx \frac{\theta_E}{T} e^{\left(\frac{\theta_E}{2T} - \frac{\theta_E}{T}\right)} \approx \frac{\theta_E}{T} e^{\left(-\frac{\theta_E}{2T}\right)}, T \rightarrow 0, e^{\left(-\frac{\theta_E}{2T}\right)} \text{ 下降的趨勢，比 } \frac{\theta_E}{T} \text{ 升高}$$

的趨勢還要快， $e^{\left(-\frac{\theta_E}{2T}\right)} \rightarrow 0, f = 0, C_{v,m} = 0$ 所以低溫下， $C_{v,m} = 3R f^2 = 3R$

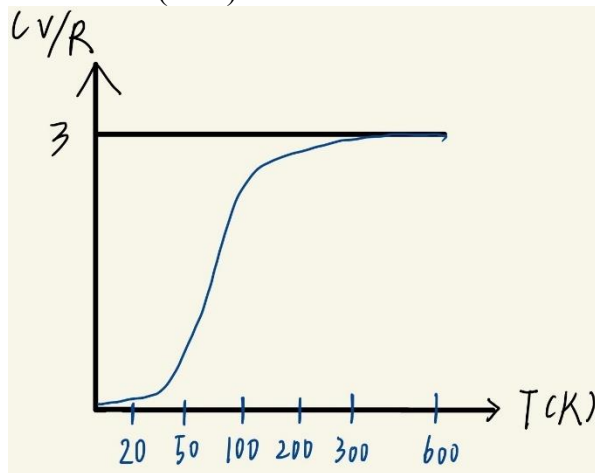
$$\times 0^2 = 0$$

高溫下， T 趨近於 ∞ ，使用 Taylor expansion 把 f 展開然後取到一階項(因為越

$$\text{高次項，} T \text{ 只會越小)，} f \approx \frac{\theta_E}{T} \frac{1 + \frac{\theta_E}{2T}}{1 + \frac{\theta_E}{T} - 1} \approx \frac{\theta_E}{T} \frac{1 + \frac{\theta_E}{2T}}{\frac{\theta_E}{T}} = \frac{\theta_E}{T} \frac{1}{\frac{\theta_E}{T}} = 1, \text{ 所以高溫下}$$

$$C_{v,m} = 3R f^2 = 3R \times 1^2 = 3R$$

(c) The copper metal has an approximate Einstein temperature of 300 K, plot a figure of heat capacity vs. temperature showing C_v/R at 20, 50, 100, 200, 300 and 600 K. (10%)



5. For an electron in the hydrogen 1s orbital: $\psi = \frac{1}{\sqrt{\pi}} \left(\frac{1}{a_0} \right)^{3/2} e^{-r/a_0}$

(a) Calculate the average value of r . (10 pts)

$$\begin{aligned} \langle r \rangle &= \int_0^\infty \psi^* \hat{r} \psi d\tau = \int_0^\infty \gamma R_{nl}^2 4\pi r^2 dr = \int_0^\infty \left(\frac{1}{\sqrt{\pi}} \left(\frac{1}{a_0} \right)^3 e^{-2r/a_0} 4\pi r^3 dr \right. \\ &= 4 \left(\frac{1}{a_0} \right)^3 \int_0^\infty r^3 e^{-2r/a_0} dr \\ &\left(\int_0^\infty r^n e^{-\beta r} dr = \frac{n!}{\beta^{n+1}} \right) \\ \langle r \rangle &= 4 \left(\frac{1}{a_0} \right)^3 \frac{3!}{\left(\frac{2}{a_0} \right)^{3+1}} = \frac{3a_0}{2} \end{aligned}$$

(b) Find the distance where the radial distribution function is a maximum (5 pts)

$$P(r)dr = \psi^2 r^2 dr = \frac{1}{\pi} \left(\frac{1}{a_0} \right)^3 r^2 e^{-2r/a_0} dr$$

$$\text{當 } \frac{P(r)}{dr} = 0 \text{ 有最大值}$$

$$\frac{d \left(\frac{1}{\pi} \left(\frac{1}{a_0} \right)^3 \int_0^\infty r^3 e^{-2r/a_0} \right)}{dr} = \frac{1}{\pi} \left(\frac{1}{a_0} \right)^3 \left(2r e^{-2r/a_0} + r^2 \left(\frac{-2r}{a_0} \right) e^{-2r/a_0} \right) = 0$$

$$\rightarrow 1 + r \left(\frac{-1}{a_0} \right) = 0, r = a_0$$

(c) Calculate the average value of the potential energy of the electron. (5 pts)

$$\begin{aligned} \langle v \rangle &= \int_0^\infty \psi^* \hat{v} \psi d\tau = \frac{1}{\pi} \left(\frac{1}{a_0} \right)^3 \int_0^\infty e^{-2r/a_0} \left(-\frac{ze^2}{4\pi\epsilon_0 a_0} \right) r^2 4\pi dr \\ &= -\frac{1}{\pi a_0^3} \frac{ze^2}{4\pi\epsilon_0} \frac{a_0^2}{4} 4\pi = -\frac{ze^2}{4\pi\epsilon_0 a_0} \end{aligned}$$